TURBULENCE IN CMEs

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Solar wind structures associa ted with coronal mass ejections have many distinguishing attributes. These attributes include for example, the relative abundance of helium, bi-direc tio nal electron streaming and high intensity helical largescale magnetic fields. Another disti nguishi ng attribute may be the character of the MHD turbulence inside C\ 4Es. We study fluct uations of the CME magnetic fields measured by ISEE3 and Ulysses spacecraft. The spectra and high-order structu re functions characterizing the intermittency are found. The results are compared with those obt ained for the quasi-station ary solar wind. The anisot ropy of fluctuations in CMEs is estimated. This anisotropy is higher than that in the interplanetary medium due to the stronger magnetic fields. The helical properties of the turbulence inside CMEs are discussed.

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